

## Electronic Structure and Metal-insulator Transition of $\text{Ca}_{2-x}\text{Sr}_x\text{RuO}_4$ Studied by NEXAFS

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**Introduction:**  $\text{Sr}_2\text{RuO}_4$  is still the only superconductor that is isostructural to a high-temperature superconductor but does not contain copper. Its superconducting and magnetic properties are unconventional [1,2], and with three Ru-O derived bands crossing  $E_F$  it exhibits a rich electronic structure [1,3,4]. Substituting Sr by Ca leads to distortions of the crystal structure and to a (Mott-type?) metal-insulator transition at  $x \approx 0.2$ ; for Ca-rich samples antiferromagnetic ordering becomes possible [5].

**Methods and Materials:** With polarization-dependent NEXAFS at the O 1s absorption edge we investigated the unoccupied electronic structure for a series of  $\text{Ca}_{2-x}\text{Sr}_x\text{RuO}_4$  single crystals, with emphasis on samples near the metal-insulator transition. Bulk-sensitive fluorescence detection was employed.

**Results and Conclusions:** Compared to pure  $\text{Sr}_2\text{RuO}_4$  (cf. [3]) the main spectral features change moderately with Ca content; a more detailed analysis is facilitated by fitting with asymmetric Gaussians. The changes of peak positions with  $x$  reflect well the symmetry changes of the  $\text{RuO}_6$  octahedra, while the development of the spectral weight of the various features points to charge carrier redistribution near  $E_F$  as well as to changes in bandwidths and hybridization strengths due to the structural distortions. The most intriguing aspect of this study is the clear Hubbard-like splitting of a near- $E_F$  feature that becomes stronger with increasing Ca content: although the Hubbard correlation energy  $U$  for Ru is not very large in itself, it is very close in value to the bandwidths of two near- $E_F$  bands [1] – and in turn our NEXAFS observation of the Ca-induced splitting directly shows the increased importance of correlation effects due to band narrowing for the Ca-rich compounds.

**References:** [1] Y. Maeno, *Physica C* **282-287**, 206 (1997) and references therein; [2] T. Imai, A.W. Hunt, K.R. Thurber, and F.C. Chou, *Phys. Rev. Lett.* **81**, 3006 (1998); [3] M. Schmidt, T.R. Cummins, M. Bürk, D.H. Lu, N. Nücker, S. Schuppler, and F. Lichtenberg, *Phys. Rev. B* **53**, R14761 (1996); [4] D.J. Singh, *Phys. Rev. B* **52**, 1358 (1995); [5] O. Friedt, M. Braden, G. André, P. Adelman, S. Nakatsuji, and Y. Maeno, [cond-mat/0007218](#); S. Nakatsuji and Y. Maeno, *Phys. Rev. Lett.* **84**, 2666 (2000); *Phys. Rev. B* **62**, 6458 (2000).